

# Diamond Scattering Detectors for Compton Telescopes

Completed Technology Project (2018 - 2020)



## Project Introduction

The objective of the proposed work is to demonstrate the suitability of artificial single-crystal diamond detectors (SCDDs) for use as the scattering medium in Compton telescopes for medium-energy gamma-ray astronomy. SCDDs offer the possibility of position and energy resolution comparable to those of silicon solid-state detectors (SSDs), combined with efficiency and timing resolution so-far only achievable using fast scintillators. When integrated with a calorimeter composed of fast inorganic scintillator, such as CeBr<sub>3</sub>, read out by silicon photomultipliers (SiPMs), SCDDs will enable a compact and efficient Compton telescope using time-of-flight (ToF) discrimination to achieve low background and high sensitivity. This detector development project will be a collaboration between the University of New Hampshire (UNH) and Southwest Research Institute (SwRI). The proposed work represents an innovative combination of detector technologies originally conceived separately for high-energy astronomy (fast scintillators read out by SiPMs; UNH) and space plasma/particle physics (SCDDs; SwRI). Recently SwRI has demonstrated that SCDDs fabricated using chemical vapor deposition (CVD) show good energy resolution ( $\sim 7$  keV FWHM), comparable to silicon SSDs, with much faster time response ( $\sim$ ns rise time) due to higher electron/hole mobilities. They are also temperature- and light-insensitive, and radiation hard. In addition, diamond is low-Z, composed entirely of carbon, but relatively high-density (3.5 g cm<sup>-3</sup>) compared to silicon or organic scintillator. SCDDs are therefore an intriguing possibility for a new Compton scattering element: if patterned with  $\sim$ mm-sized readout electrodes and combined with a fast inorganic scintillator calorimeter, SCDDs could enable a compact but efficient Compton telescope with superior angular and energy resolution, while maintaining ToF background rejection. Such an instrument offers the exciting potential for unprecedented sensitivity, especially at energies  $< 1 - 2$  MeV, on a small-scale mission utilizing recently available SmallSat buses (payload mass  $< 100$  kg). We propose to demonstrate this by constructing and testing a small proof-of-concept prototype and, based on its performance, using Monte Carlo simulations to explore the possibilities of furthering MeV science using relatively small-scale space missions.



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## Organizational Responsibility

### Responsible Mission Directorate:

Science Mission Directorate (SMD)

### Lead Organization:

University of New Hampshire-  
Main Campus

### Responsible Program:

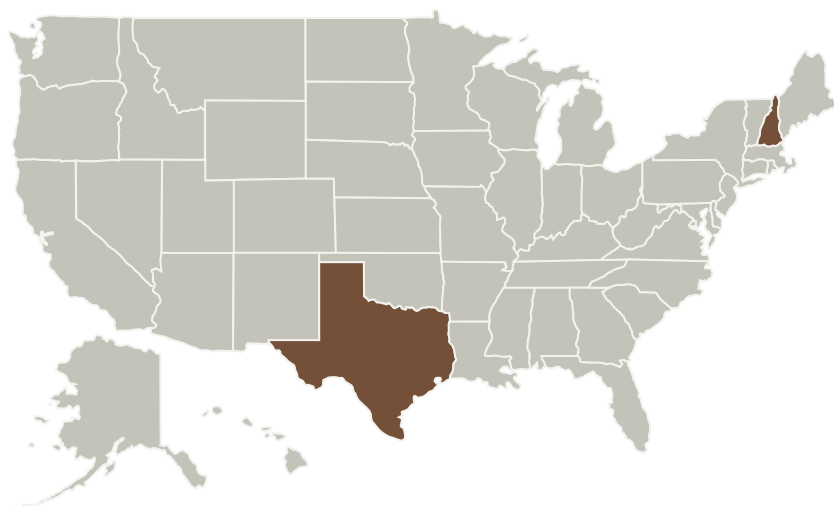
Astrophysics Research and  
Analysis

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## Primary U.S. Work Locations and Key Partners



Organizations Performing Work	Role	Type	Location
University of New Hampshire-Main Campus	Lead Organization	Academia	Durham, New Hampshire
Office of Sponsored Research and Programs - NC Central University	Supporting Organization	Academia	Durham, New Hampshire
Southwest Research Institute - San Antonio(SWRI)	Supporting Organization	Academia	San Antonio, Texas

Primary U.S. Work Locations	
New Hampshire	Texas

## Project Management

**Program Director:**

Michael A Garcia

**Program Manager:**

Dominic J Benford

**Principal Investigator:**

Peter F Bloser

**Co-Investigators:**

Lisa Scigliano

Mark Mcconnell

Keiichi Ogasawara

James M Ryan

## Technology Areas

**Primary:**

- TX08 Sensors and Instruments
  - └ TX08.1 Remote Sensing Instruments/Sensors
  - └ TX08.1.1 Detectors and Focal Planes

## Target Destination

Outside the Solar System